

# Propagation of Ultrashort, Intense Laser Pulses Through the Atmosphere



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### What is an ultrashort laser pulse?

### laser parameters

peak power	GW to PW
intensity	10 <sup>12</sup> to 10 <sup>23</sup> W/cm <sup>2</sup>
pulse duration	10 <sup>-14</sup> to 10 <sup>-12</sup> seconds
average power	1 W- 1 kW

### comparison of electric field strengths

- laser electric field:  $E_{laser}$ [V/cm] = 27.5 I<sup>1/2</sup> [W/cm<sup>2</sup>]  $E_{laser} \sim 10^8$  to  $10^{13}$  V/cm
- atomic electric field (over the barrier ionization):  $E_{atom} \sim q/4\pi\epsilon_0 a^2 \sim 5x10^9 \text{ V/cm}$
- Schwinger field (potential difference of a rest mass across a Compton wavelength):

$$E_{sch} = m_e^2 c^3 / q\hbar \approx 10^{16} \text{ V/cm}$$

### Fundamentally different interactions than long pulses

### laser wakefield acceleration of electrons

a laser pulse propagating through plasma ponderomotively drives an electrostatic wave that can accelerate electrons to MeV-GeV energies

### radiation pressure acceleration of ions

a laser pulse incident on an overdense target acts as a light sail/piston, ponderomotively accelerating ions

positron and gamma ray sources

a laser pulse incident on a counterpropagating multi-GeV electron beam yields gamma rays that pairproduce

generation of ultrabroadband radiation

a laser pulse propagating through a nonlinear medium undergoes octave-spanning spectral broadening and generation of far out-of-band radiation



W. Lu et al., PRSTAB, 10 (2007).



S. V. Bulanov et al., Phys. Plasmas, 17 (2010).



extreme light Infrastructure website



S. Varma, thesis defense (2011).



- Much of our understanding of optical propagation through atmospheric turbulence is based on the medium's **linear response** to the optical field.
- Relatively few studies of turbulence and nonlinear filamentation and nonlinear self-focusing\*. The majority of these studies investigated cases where the peak laser power is larger than the self-focusing power of air (filamentation regime).
  - **Kandidov** et al.: modeling of nonlinear self-focusing (NLSF) in turbulence using phase screens. [Quantum Electron. 29, 911 (1999)]
  - Chin et al.: characterized filament wander in turbulence [Appl. Phys. B 74, 67–76 (2002)]
  - **Salame** et al.: laboratory experiments over regions of extended turbulence [Appl. Phys. Lett. 91, 171106 (2007)]
  - **Houard** et al.: characterized the competition between self-focusing and modulational instability, investigate filament stability in turbulence [Phys. Rev. A 78, 033804 (2008)]
  - Penano et al.: calculation of filamentation onset/suppression in turbulence [JOSA B 31, 963 (2014)]
- In this study, we present a way to propagate long ranges through strong atmospheric turbulence that avoids the "filamentation" regime.

# **Characterization of Atmospheric Turbulence**

- turbulence: random fluctuations of air currents in the form of eddies driven by large scale thermal gradients
- Eddy scale distribution is described by the modified von Karman spectrum:

$$\Phi_n(k) = .033C_n^2 \frac{e^{-\ell_0^2 k^2}}{(k^2 + L_0^{-2})^{11/6}}$$

$$L_0$$
: outer scale (1 – 100 m)  
 $\ell_0$ : inner scale (~1-10 mm)



 $C_n^2$ : refractive index structure constant (meteorological strength)

- Rytov variance parameterizes the optical effect:  $\sigma_R^2 = 1.23 C_n^2 k^{7/6} L^{11/6}$  $\sigma_R^2 << 1$  (optically weak) ;  $\sigma_R^2 > 1$  (optically strong)
- Transverse coherence length of a plane wave:  $\rho_0 = (1.46k^2C_n^2L)^{-3/5}$



### Turbulence and its effect on laser propagation



 $(\sim 1 \text{ um wavelength})$ 

- Median turbulence over near-surface, horizontal path\*:  $C_n^2 = 3 \times 10^{-15} \text{m}^{-2/3}$ \*C.I. Moore, et al., Proc. of SPIE Vol. 5892(SPIE, Bellingham, WA, 2005):
- Strong turbulence regime can be realized within several kilometers:  $\begin{bmatrix} 5 \text{ km path} \\ \sigma_R^2 = 3.2 \end{bmatrix}$



#### Turbulence refractive index



# Self-channeling through turbulence



Turbulence effects:

spreading is due to small scale turbulence wander is due to large scale turbulence



*R* is the beam spot size

- Self-channeling concept
  - Use small spot size (smaller than the coherence length or inner scale) to avoid spreading by turbulence
  - Propagate near the critical power to cancel diffractive spreading



Channeling can persist for many Rayleigh lengths

Simulations Indicate Effectiveness at Long Range ( $z >> Z_R$ )



 Simulations solve NLSE in (x,y,z,t) including GVD, extinction, and turbulence

$$\left[2ik\left(\frac{\partial}{\partial z}+\alpha\right)-k\beta_2\frac{\partial^2}{\partial\tau^2}+\nabla_{\perp}^2\right]A = -2k_0^2\left(\delta n_T + n_2I\right)A$$

100+ independent realizations of turbulence for a given  $C_n^2$ 



• Characterize channeling by how much energy remains within the initial beam radius, W<sub>0</sub>.

J. Peñano, J.P. Palastro, B. Hazi, M.H. Helle, and G.P. DiComo. Self-channeling of high-power läser pulses through strong atmospheric turbulence. *Phys. Rev. A*, 96(1):013829, 2017.

 $z = 15 Z_R$ 



# Channeling is dependent on relationship between spot size (W0), inner scale (I0), and coherence length (r0)



- (i)  $W_0 < (\ell_0, \rho_0)$  long-range channeling (> 10 Z<sub>R0</sub>)
- (ii)  $\ell_0 < W_0 < \rho_0$  rms spot increases, central hot-spot remains intact
- (iii)  $W_0 >> (\ell_0, \rho_0)$  nonlinear self-focusing is not effective



## Channeling is degraded by field incoherence



- Channeling of the central hot spot is lost when the coherence length of the field becomes smaller than the beam rms radius. (beam and "reservoir" become incoherent)
- Coherence length of the channeling Gaussian beam is characteristic of that of a plane wave.

Solid curves: theoretical plane wave coherence radius for  $---- W_0/\ell_0 = 0.1$  $----- W_0/\ell_0 = 10$ 



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### **Longitudinal Compression of Laser Pulses in Air**

- Air is a dispersive medium at optical frequencies, i.e., frequencies have different group velocities
- Ultrashort pulses have sufficient bandwidth such that atmospheric dispersion can compress or stretch the pulse significantly over hundreds of meters of propagation
- Pulse duration after distance z:

$$T(z) = T_o \left[ \left( 1 + \beta_o \frac{z}{Z_T} \right)^2 + \left( \frac{z}{Z_T} \right)^2 \right]^{1/2}$$
$$Z_T = T_o^2 / 2 \left| \beta_2 \right| \qquad \beta_0 = \frac{T_0}{2\beta_2} \frac{\partial T}{\partial z} \Big|_{z=0}$$

 Arranging spectral content of the pulse so frequency decreases from front to back (negative chirp) results in longitudinal compression (for wavelengths ~ 1μm where β<sub>2</sub>(ω)>0)





# In the presence of extinction, the pulse must be chirped to preserve peak power and maintain channeling

- NRL PPD
- Pulse duration vs. z:  $T(z) = T_0 \left[ (1 + \beta_0 z/Z_T)^2 + (z/Z_T)^2 \right]^{1/2}$ ;  $Z_T = T_0^2/(2|\beta_2|)$
- Optimal chirp:  $\beta_0 \approx \alpha T_0 E_0 / (\beta_2 P_{NL}) \longrightarrow T'(z) = E'(z) / P_{NL}^{-1}$



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2015-16 Lab environment at NRL kTFL laser (5 mJ, 35 fsec, 1 kHz, 800 nm)

Nonlinear channeling in turbulence

#### **AFRL Facility** 2016 Pheenix laser (40 TW, 35 fsec, 10 Hz, 800nm)

Nonlinear channeling in deep turbulence (Rytov > 5)

#### **Carderock Facility** 2017

Astrella laser (7 mJ, 35 fsec, 1 kHz, 800 nm) Nonlinear propagation through turbulence Pulse compression Control of nonlinear focal range

### Validation Experiments at Increasing Distances









### Laboratory Turbulence Generator for high-power laser propagation experiments\*





- 5 orders of magnitude range of turbulence
- Kolmogorov spectrum over the inertial sub-range
- Ideal for high-power experiments (no phase plates to damage)



\*G. DiComo et al. "Implementation of a long range, distributed-volume, continuously variable turbulence generator," Applied Optics 55, 5192 (2016). S. Pond, R. W. Stewart, and R. W. Burling, Journal of Atmospheric Sciences, vol. 20, pp. 319-324 (1963).

## **Experimental Facility for Validation at ~850m**





- Continuous indoor path with low ambient turbulence
- Distributed turbulence with tunable Rytov from  $10^{-3}$  to >10
- Two weeks to build



- Laser operation during graveyard shift over 2 months
- Laser intensity profile and pulse length fully characterized

## **Portable USPL for propagation experiments**



### Coherent ASTRELLA

- Wavelength: 800nm
- Rep Rate: 1 kHz
- Energy: 7 mJ
- Pulse length: 35fs
- Beam Quality:  $M^2 < 1.25$
- Integrated Vitara seed laser, revolution pump laser, STAR regenerative amplifier and sealed, compact stretcher/compressor
- HASS\* verified (ruggedized for portability)
- All major sub-systems thermally-stabilized for reliable long-term performance
- Water-only cooled Ti:Sapphire rod assembly for improved beam quality and thermal management ( $M^2 < 1.25$ , stability <0.5% rms)
- Sealed stretcher/compressor (pulse width <35 fs to 3ps)



#### Laser housing open for setup

Typical Near Field Mode Quality



### **Self-Channeling Observed in Strong Turbulence**





Weak Turbulence ( $\sigma_{\rm R} \sim 0.05$ )



Strong Turbulence ( $\sigma_{\rm R} \sim 5$ )



High-power beam **remains small (<< diffraction limit)** in strong turbulence

## **Experimental measurement of temporal profile**



- Built SHG FROG to measure complex field (Can measure low intensities)
- Pulse was characterized at laser and received end of range
- Nonlinear effects such as self-phase modulation and self-steepening can be observed for the high power pulse

### Frequency-Resolved Optical Gating (FROG)



SHG FROG is the most sensitive version of FROG.



# Experiments validate nonlinear self-focusing models in strong turbulence





### Beam wander is consistent with linear physics







- For a given rms spot size, Gaussian beams have the longest diffraction length compared with any other beam shape
- Experimentally generated laser beams are never purely Gaussian, nor perfectly coherent
- What are the nonlinear self-focusing properties of non-Gaussian, partially coherent beams? Do they afford any propagation advantages?
  - Bessel beams
  - Partially coherent beams



• The ensemble averaged beam spot size, w(z), along the propagation axis z is

$$w^{2}(z) = \frac{2}{P} \int r^{2} \langle I(r,z) \rangle d^{2}r - 2 \langle R_{c}^{2}(z) \rangle$$

• The spot size satisfies the following **mathematical** identity

$$\frac{\partial^2 w}{\partial z^2} = \frac{\lambda_0^2}{n_0^2 \pi^2} \frac{M^4(z)}{w^3}$$

where the ensemble beam quality (spreading angle normalized to the the Gaussian spreading angle) is:

$$M^{4}(0) = \frac{n_{0}^{2}\pi^{2}}{\lambda_{0}} \left\{ -\frac{1}{4} \left( \left. \frac{\partial w^{2}}{\partial z} \right|_{0} \right)^{2} + \frac{w(0)^{2}}{2} \left. \frac{\partial^{2} w^{2}}{\partial z^{2}} \right|_{0} \right\}$$

- In the absence of turbulence, it can be shown that M<sup>4</sup> is a conserved quantity. If M<sup>4</sup> can be evaluated at z=0 then we will know how the ensemble of beams will propagate.
- Define the critical power as the condition  $M^2 = 0$

The Critical Power of an Arbitrary Beam Shape (cont.)



$$M^{4}(0) = \frac{n_{0}^{2}\pi^{2}}{\lambda_{0}} \left\{ -\frac{1}{4} \left( \left. \frac{\partial w^{2}}{\partial z} \right|_{0} \right)^{2} + \frac{w(0)^{2}}{2} \left. \frac{\partial^{2}w^{2}}{\partial z^{2}} \right|_{0} \right\}$$

Only the 2<sup>nd</sup> term is important in evaluating M<sup>4</sup> (the first term represents a "lens" contribution which does not affect M<sup>4</sup>):

$$\frac{\partial^2 w^2}{\partial z^2} = \frac{2}{P} \int r^2 \left\langle \frac{\partial^2 I}{\partial z^2} \right\rangle d^2 r = \frac{2}{P} \frac{c n_0}{8\pi} \int r^2 \left\langle \left| \frac{\partial A}{\partial z} \right|^2 + A^* \frac{\partial^2 A}{\partial z^2} + \text{c.c.} \right\rangle d^2 r$$

• The physics of the laser propagation is introduced through the nonlinear Schodinger equation (NLSE)

$$\frac{\partial A}{\partial z} = \frac{i}{2k_0} \left( \nabla_{\perp}^2 A + 2k_0^2 \frac{n_0^2 n_2 c}{8\pi} |A|^2 A \right)$$
  
diffraction self-focusing

• Using a Gaussian beam as the initial condition in this formalism recovers the usual expression for the self-focusing power:

$$P_G = \lambda^2 / (2\pi n_0 n_2)$$







• The Gaussian-Schell model is a way to generate a partially-coherent field

$$A(\vec{r}) = \left(\frac{8\pi I_0}{cn_0}\right) \left(\int a(\vec{k})e^{i\vec{k}r}d^3k\right) e^{-r^2/R_0^2}$$

where *a(k)*, the angular correlation function of the field, and the intensity are  $\langle a^{\star}(\vec{k})a(\vec{k'})\rangle = \frac{\rho_c^2}{\pi}e^{-k^2\rho_c^2}\delta(\vec{k'}-\vec{k}) \qquad \qquad I(\vec{r}) = \frac{cn_0}{8\pi}\langle A^{\star}(\vec{r})A(\vec{r})\rangle$ 

• By applying the formalism described in the previous slides, we obtain for the normalized spreading angle

$$\begin{split} M^4(0) &= n_0^2 \left( 1 + \frac{R_0^2}{2\rho_c^2} - \frac{P}{P_{\mathrm{crit}}} \right) \qquad P_{crit} = \frac{\lambda_0^2}{2\pi n_0 n_2} \\ \uparrow & \uparrow & \mathsf{focusing} \\ \mathsf{defocusing} \end{split}$$

 The defocusing term due to incoherence is identical in form for propagation in turbulence with the coherence length replaced by the Fried parameter (the atmospheric coherence length)



m = 0n = 1

Field envelope: 
$$A = \left(\frac{8\pi I_0}{n_0 c}\right)^{\frac{1}{2}} J_m(rj_{mn}/R_0) \exp(im\theta)(1 - H(r - R_0))$$
;  $j_{mn}$  is the *n*<sup>th</sup> root of  $J_m$  aperture radius:  $R_a$ 

Power of a Bessel beam:  $P_{mn} = \pi R_0^2 I_0 J_{m-1}^2(j_{mn})$ 

#### **Comparison with apertured Gaussian beam:**

- For both beams to have equivalent RMS radius:  $\frac{W_0^2}{2} \frac{R_0^2 \exp(-2R_0^2/W_0^2)}{1 \exp(-2R_0^2/W_0^2)} = \langle r^2 \rangle_{mn}$
- For apertured Gaussian with intensity  $I_G$  to have equivalent power:

$$\frac{I_G}{I_B} = \frac{2R_0^2}{W_0^2} \frac{J_{m-1}^2(j_{mn})}{1 - \exp(-2R_0^2/W_0^2)}$$

Figures show intensity profiles of Bessel beams (solid curves) and apertured Gaussian beams (dashed curves) with equivalent RMS spot size and power.



0.8

 $I/I_G$ 

0.2

0.0

0.2

0.4

0.6

0.8



Self-focusing power ( $P_B$ ) of a Bessel beam transmitted through an aperture of radius  $R_0$ 

$$P_B = \frac{P_G}{2} j_{mn}^2 J_{m-1}^2(j_{mn}) \frac{\int_0^{j_{mn}} dx \left[\frac{x}{4} \left(J_{m-1}(x) - J_{m+1}(x)\right)^2 + \frac{m^2}{x} J_m^2(x)\right]}{\int_0^{j_{mn}} dx x J_m^4(x)}$$

 $P_G = \lambda^2/(2\pi n_0 n_2)$  is the self-focusing power of a Gaussian beam



### Bessel beams are very unstable to filamentation



- For Bessel beams at the "critical power", filamentation is governed by smaller-scale filamentation.
- For example, the m=0 mode has a filamentation distance given by:

$$L_{B1} \approx 0.257 (\pi r_1^2 / \lambda) / \sqrt{P_{B1} / P_G - 1}$$

which can be much shorter than the Rayleigh length of the whole beam.





- A gas or plasma can be used to create a radial refractive index variation that can act as a lens with very high damage threshold for high-intensity laser pulses.
- NRL recently created a vortex flow gas lens that acts as a negative axicon, and can be used in conjunction with a focusing lens to create Bessel beams.

#### Lens schematic



**3D printed lens** 





Formation of annular beam (focuses to a Bessel beam)

### NRL's newly-renovated ultrashort pulse laser facility



- 350 sq. ft. ISO 4 cleanroom
- 40TW 5Hz high energy laser
- 0.5TW 1kHz repetition rate laser
- Tunable OPA
  - Currently: 1.1-2.6um
  - 200nm-15um July 2018
- Laser-driven electron and ion accelerators



A unique facility that enables high-precision, pumpprobe experiments involving any combination of photons at almost *arbitrary wavelengths* (THz thru Xray) and high energy particles (electrons and ions)

- Ultrafast laser-matter interactions and propagation
- Novel radiation sources





- Nonlinear self-channeling is achieved when the laser pulse power is close to the self-focusing power and the transverse dimensions are smaller than the turbulence coherence scale.
- Experiments demonstrate
  - Self-channeling over many Rayleigh lengths achieved
  - Self-channeling in deep turbulence
  - Self-channeling maintained over distances where temporal spreading (GVD) becomes important
  - quantitative validation of simulations
- Future work to include propagation of partially coherent, intense beams in atmospheric turbulence



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